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March 25, 2019







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Optical telescope



Radio telescope



- Square Kilometer Array
- Frequency range of [50; 350]MHz
- 130000 antennas spread between 500 stations
- Compared to the best similar instrument :
 - 25% better resolution
 - 8× more sensitive
 - $135 \times$ the survey speed



A first station overview



 $N_a = 256$ antennas irregularly arranged. The Aperture Array Verification System 1 (AAVS1).

Sources : https://www.skatelescope.org

An example of a wired antenna

Sketch of the running of antenna by converting a sine electric current into a EM wave.

Field regions



Calibration of the SKA-low antenna array using drones Radiation Pattern

$$\mathbf{F}(\theta,\varphi) = \lim_{\rho \to \infty} \frac{\mathbf{E}(\rho,\theta,\varphi)}{\max_{\theta,\varphi} \mathbf{E}(\rho,\theta,\varphi)} = F_{\nu} \, \mathbf{e}_{\theta} + F_{h} \, \mathbf{e}_{\varphi}$$

Embedded Element Pattern (EEP)

Radiation pattern when an antenna *i* is on with the other passively terminated.





Sources : T. Zwick - Antennen und Mehrantennensysteme (KIT)

Flight strategies











Calibration procedure



Calibration methods



Formulation as N_a convex optimization problems

$$\min_{\mathbf{c}' \in \mathbb{C}^{N_a}} ||\mathbf{F}_{\mathbf{v}}^i - \bar{\mathbf{F}}_{\mathbf{v}}^i||_2^2 + ||\mathbf{F}_h^i - \bar{\mathbf{F}}_h^i||_2^2 \qquad \forall i = 1 \dots N_a$$

This is a least-square problem and its optimal solution satisfies

$$\begin{pmatrix} \mathbf{B}_{h}^{i} \\ \mathbf{B}_{v}^{i} \end{pmatrix}^{\mathrm{H}} \underbrace{\begin{pmatrix} \mathbf{B}_{h}^{i} \\ \mathbf{B}_{v}^{i} \end{pmatrix}}_{\mathbb{C}^{2 N_{\mathsf{mes}} \times N_{\mathsf{a}}} \underbrace{\mathbf{c}^{i}}_{\mathbb{C}^{N_{\mathsf{a}}}} = \begin{pmatrix} \mathbf{B}_{h}^{i} \\ \mathbf{B}_{v}^{i} \end{pmatrix}^{\mathrm{H}} \underbrace{\begin{pmatrix} \mathbf{F}_{h}^{i} \\ \mathbf{F}_{v}^{i} \end{pmatrix}}_{\mathbb{C}^{N_{\mathsf{a}}}} \qquad \forall i = 1 \dots N_{\mathsf{a}}$$

Error definition in dB

$$e_{\theta,\varphi} = 10\log_{10}\left(\left|\mathbf{F}_{v} - \bar{\mathbf{F}}_{v}\right|^{2} + \left|\mathbf{F}_{h} - \bar{\mathbf{F}}_{h}\right|^{2}\right) - 10\log_{10}\max_{\theta,\varphi}\left\{|\mathbf{F}_{v}|^{2} + |\mathbf{F}_{h}|^{2}\right\}$$

Metrics

 $\begin{array}{ll} \mathsf{Maximum error} & e_{\mathcal{M}} \triangleq \max_{\theta,\varphi} e_{\theta,\varphi} \\ \mathsf{Mean error} & \mu_e \triangleq \mathbb{E}\{e_{\theta,\varphi}\} \end{array}$

Results

Results



Far-field calibration

No flight restriction

- $h \triangleq 1 \text{ km}$ height to reach Fraunhofer distance
- $a = 1.7 \, \text{km}$

| Flight strategy | e_M (dB) | $\mu_{e} (dB)$ |
|-----------------|------------|----------------|
| Spiral | -60.9 | -74 |
| Grid | -63.4 | -75.5 |
| Cuts | -29.8 | -45.3 |
| Random | -60 | -74.3 |
| Equi-spaced | -64.3 | -76.5 |

Good behavior for all strategies except cuts



Far-field calibration

200 m flight restriction

• *h* = 1 km

• *a* = 200 m

Small error along the drone path Large error anywhere else



Near-field calibration

• *h* = 10 m

| Flight strategy | <i>e_M</i> (dB) | $\mu_e~(dB)$ |
|-----------------|---------------------------|--------------|
| Spiral | -31.1 | -41.3 |
| Grid | -31.72 | -42.5 |
| Cuts | -20.6 | -31.2 |
| Random | -28.7 | -41.0 |
| Equi-spaced | -29.7 | -41.1 |
| | | |



Better than restricted FF but worse than unrestricted FF

Near Field



Near field

Drone attitude

| Max error (dB) | Mean error (dB) | σ_y | σ_p | σ_r | (degrees) |
|-----------------|-----------------|------------|------------|------------|-----------|
| - 48 .14 | -57.38 | 2 | 2 | 2 | |
| - 48 .46 | -59.62 | 2 | 0 | 0 | |
| -54.32 | -65.46 | 0 | 2 | 0 | |
| -54.32 | -65.14 | 0 | 0 | 2 | |



Smaller impact than position noise Error mainly due to yaw angle

Near field



Linear relationship between error and noise-level in dB scale When combined, the error is dominated by the more significant noise

Regularization

Regularization

Mathematical formulation

$$\min_{\mathbf{x} \in \mathbb{R}^n} ||A\mathbf{x} - (\mathbf{b} + \epsilon \mathbf{f})||_2^2$$

with $\mathbf{A} \in \mathbb{R}^{m \times n}$, $\mathbf{b} \in \mathbb{R}^n$ the data matrix and $\epsilon \mathbf{f}$ the perturbation vector.

A numerical example

Let

$$A = \begin{pmatrix} 0.16 & 0.10 \\ 0.17 & 0.11 \\ 2.02 & 1.29 \end{pmatrix}, \quad \mathbf{x}^{\star} = \begin{pmatrix} 1 \\ 1 \end{pmatrix}, \quad \mathbf{b} = A\mathbf{x}^{\star}$$

with $\mathbf{f} = \begin{pmatrix} 1 \\ -3 \\ 2 \end{pmatrix}$ a perturbation with $\epsilon = 0.01$.

Moore-Penrose solution is $\mathbf{x}_{\epsilon} = \begin{pmatrix} 7.0 & -8.3 \end{pmatrix}^{T}$.

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Regularization Regularization type

Tikhonov Truncated SVD (TSVD) Least square minimization with a quadratic inequality constraint (LSQI) Parameter choice strategy

Discrepancy principle Generalized cross-validation (GCD) Quasi-optimality criterion (QO)

Application on SKA-low calibration 40 20 $\mu_{\rm e} \, [{\rm dB}]$ 0 -20 -40 -60 Spiral Grid Cut Random Equi-spaced NF-Cut Non-regularized Tikhonov - GCV Tikhonov - Quasi-optimality TSVD - GCV TSVD - Quasi-optimality LSQI - Discrepancy Principle

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Regularization Regularization type

Tikhonov

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Conclusion and future works

| | | | Non Regularized regularized | | | | | | |
|------|--------------|--|---------------------------------|----------|---------|--------------|-------------|---------|--|
| | | | | Tikh-GCV | Tikh-QO | TSVD- GCV | TSVD- QO | LSQI-DP | |
| | | Spiral | 56 | -27 | -35 | -11 | -40 | -45 | |
| | ctec | Grid | 44 | -43 | -34 | -42 | -39 | -8 | |
| tric | Cut | 88 | -6 | -34 | -21 | -39 | 31 | | |
| | Res | Random | 63 | -42 | -34 | -41 | -39 | -2 | |
| FF | | Equi- spaced | 24 | -43 | -35 | -43 | -39 | -43 | |
| | Unrestricted | Spiral Grid Cut Random Equi- spaced | -74 -75 -45 -74 -76 | | | | | | |
| NF | | Spiral Grid Cut Random Equi- spaced | -41 -43 -31 -41 -41 | -44 | -33 | -43 | -40 | -43 | |

| | | | Non regu- larized | | | | | | |
|----|--------------|--|---------------------------------|--------------------------------|---------------------------------|---------------------------------|---------------------------------|------------------------------|--|
| | | | | Tikh-GCV | Tikh-QO | TSVD- GCV | TSVD- QO | LSQI-DP | |
| FF | Restricted | Spiral Grid Cut Random Equi- spaced | 56 44 88 63 24 | -27 -43 -6 -42 -43 | -35 -34 -34 -34 -35 | -11 -42 -21 -41 -43 | -40 -39 -39 -39 -39 | -45 -8 31 -2 -43 | |
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Mean Error μ_e with 256 experiments

| | | | Non regu- larized | | | | | | |
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Mean Error μ_e with 256 experiments

Going further

Extending the model

- Adding other noise sources (e.g. during the transmission)
- Taking finite ground plane effects into account

Generalization to $N_a \neq 256$ antennas

• FF case :
$$N_e \approx \max\{\underbrace{50}_{\text{Sampling the pattern}}, \underbrace{N_a}_{\text{N}_a}\}$$

• NF case : Every experiment yields N_a "independent" measurements \Rightarrow Same behavior for all N_a







Why using an array of antennas ?

- Obtaining a better angular resolution
- Increasing the sensitivity

Interferometry : Emulation of a single larger dish

